

STELLA-1.2

Quickstart Guide



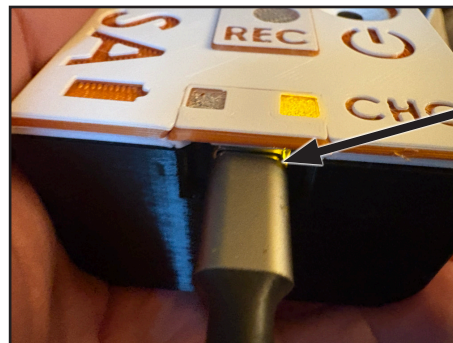
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STELLA-1.2 Quickstart

STELLA (Science and Technology Education for Land/Life Assessment) is a force multiplier for NASA Missions and programs—it amplifies their impact by breaking down barriers that make satellite remote sensing seem inaccessible, transforming mission data from abstract numbers into tangible participatory science that develops the technical workforce Earth observation systems depend upon while creating thousands of mission-literate professionals who understand why NASA investments matter.

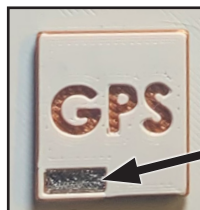
Charging



USB-C Port

Plug in USB-C cable from power to USB-C port to charge

Indicators



GPS Indicator

- Solid Red if GPS Fix
- Red Pulse if no GPS Fix



Battery Indicator

- Red flash when battery charge is low

Record Indicator

- Amber during boot sequence
- Green pulse per measurement
- Solid Red if no SD card

Charging Indicator

- Amber during charging

SD Card

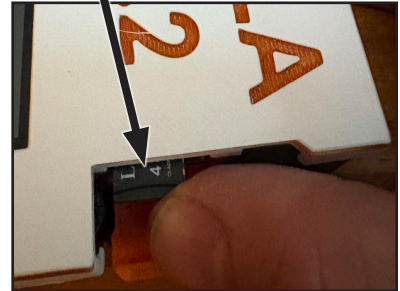
SD Card Retainer
Swing Door Closed



SD Card Retainer
Swing Door Open



SD Card
Push in to lock/release



Holders

Strap Loops

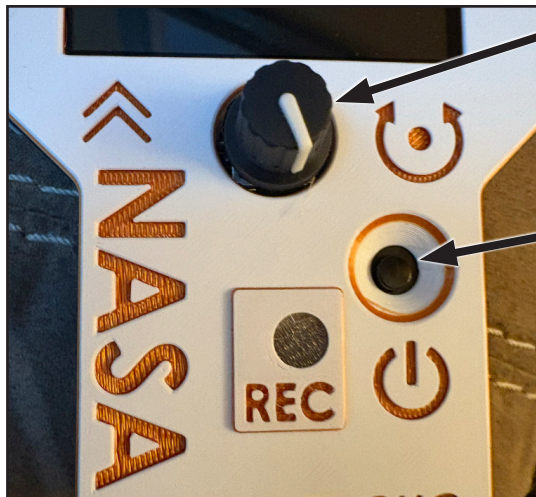


Universal Screw Mount



*Note that STELLA-1.2 is made out of PLA and should not be exposed to high temperatures for extended periods of time as it will warp. For example - do not leave in car with outside temperatures near 100°F/37.7°C for more than a few hours,

Navigation Controls



Rotary Encoder

The primary navigation device for STELLA-1.2



Turn Encoder:

The primary navigation device for STELLA-1.2



Press Encoder:

Select highlighted item and open that screen

Power Button



Press to Turn On/Off

Control Bar

Batch Number:

Click to increment, auto-increments on restart (same day), resets on day change or SD removal

Recording (Red Circle):

Click to pause

Burst Button:

Takes set measurements then stops (default: 2)

Battery %:

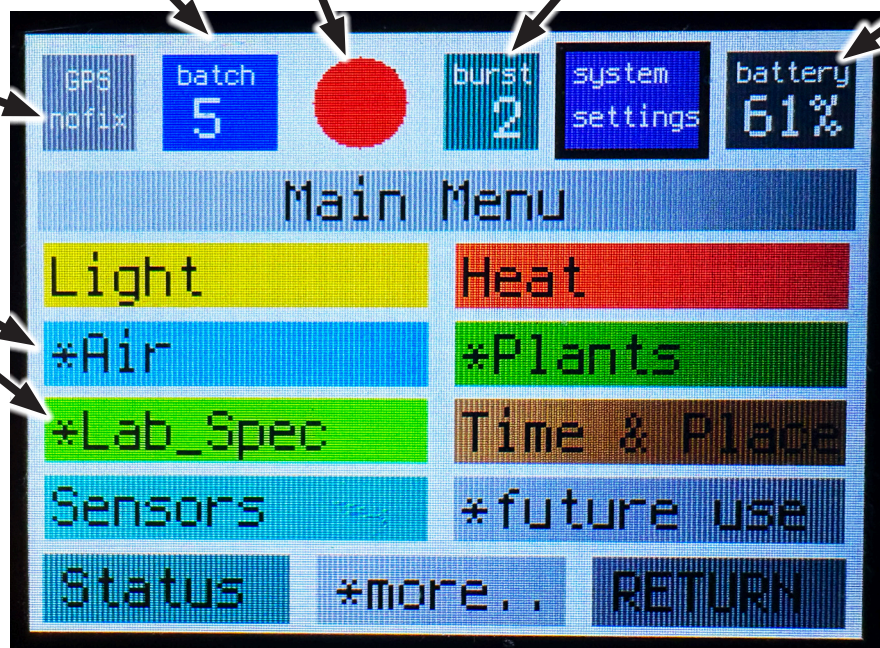
Click for Status Page

GPS Square:

Green when locked

* (asterisk):

Page is not ready yet



System Settings

Sample Interval:

Time between consecutive measurements during continuous recording.

Burst Count:

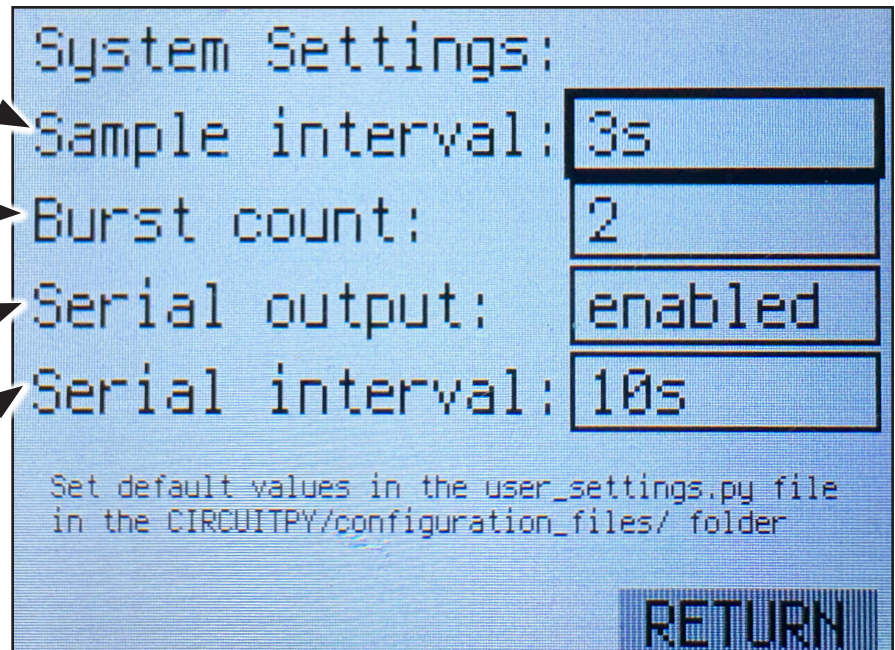
Number of measurements taken when burst button is pressed.

Serial Output:

Control data transmission to connected computer via USB

Serial Interval:

Rate of data transmission to computer (independent of SD card recording)



Status

Access: Click battery % on control bar or from main menu

Processor ID:

Last 5 digits of unique 100-digit factory ID. Identifies individual unit.

Battery Status:

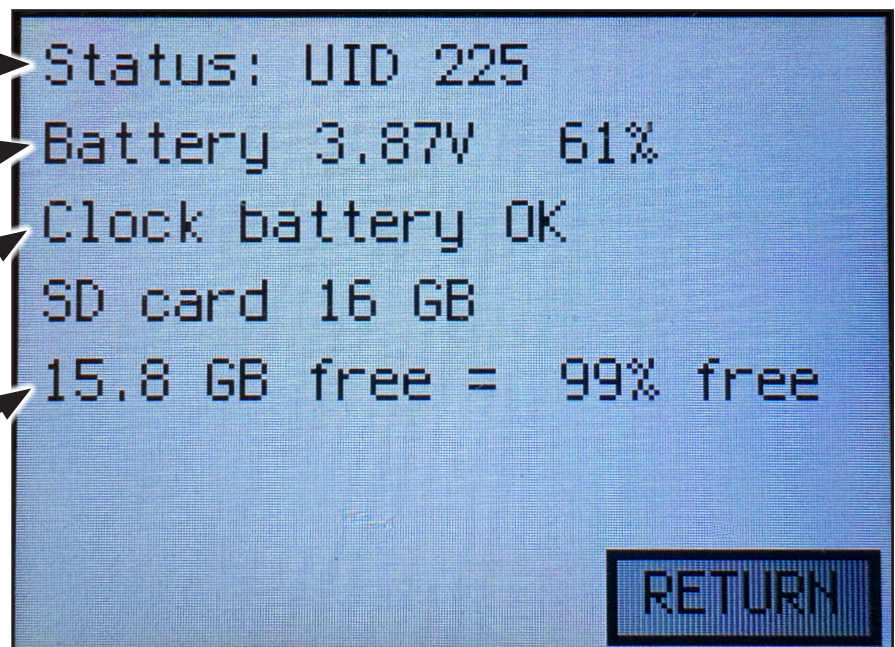
Voltage (actual) and State of Charge (%). Lookup table converts voltage to percentage.

Clock Battery:

Status: "OK" or "Low". Critical for timekeeping.

SD Card:

Total capacity and available space.



Sensors (example pages)

| supply_5V : tps61023 | |
|----------------------|-------|
| counts | 50660 |
| voltage | 5.1 |
| RETURN | |

Sensor Type and Sensor:

- 5V supply
- Counts
- Voltage
- Counts refers to the raw analog-to-digital converter (ADC) reading from the analog pin monitoring the 5V supply
- This value ranges from 0 to 65535 (16-bit ADC resolution: $2^{16} - 1 = 65536$ possible values)
- Represents the digitized voltage reading before conversion to actual voltage

| air : hdc3022 | |
|---------------|------|
| temperature_C | 19.1 |
| humidity_pct | 56.2 |
| RETURN | |

Sensor Type and Sensor:

- Air temperature
- Humidity percentage

Plugin Required

| 410nm_channel : as7265x | |
|-------------------------|-----|
| wavelength_nm | 410 |
| gain | 16 |
| int_time_ms | 354 |
| raw_counts | 0 |
| normal_ct_per_s | 0.0 |
| irrad_W_per_m2 | 0.0 |
| bandwidth_nm | 20 |
| RETURN | |

Sensor Type and Sensor:

- Wavelength nanometers
- Gain
- Integration time milliseconds
- Raw counts
- Normal counts per second
- Irradiance Watts per meter squared
- Bandwidth nanometers

Plugin Required

Time and Place

System Time:

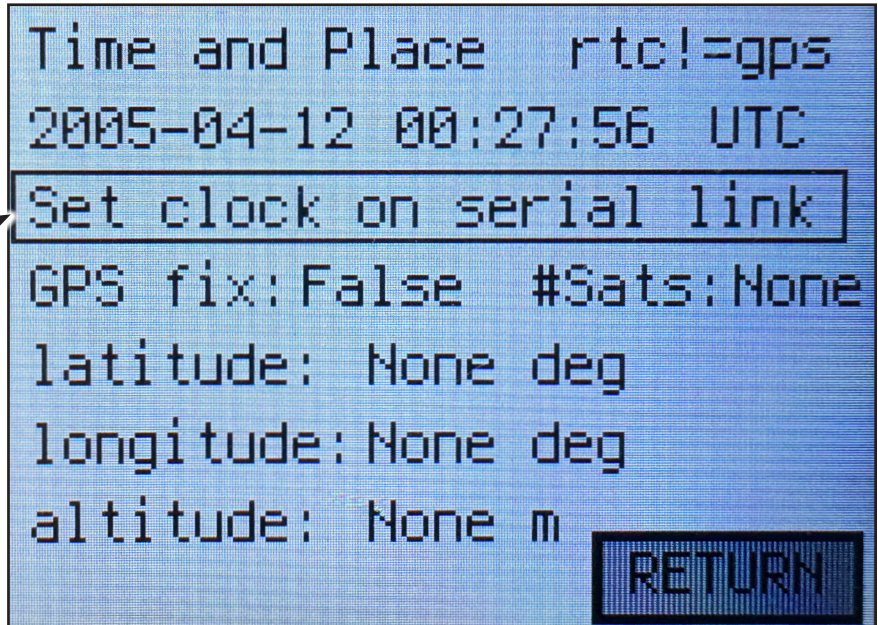
- Shows date and time from onboard clock
- Accuracy: ± 2 -3 sec/day drift
- GPS sync: checks once/min, auto-sets time
- Shows "RTC != GPS" when mismatch

Manual Time Setting:

"Set Clock" button for indoor use via serial link. GPS overrides when fix acquired.

GPS Information:

- Coordinates: Lat, Lon, Altitude
- "Fix = True" when locked

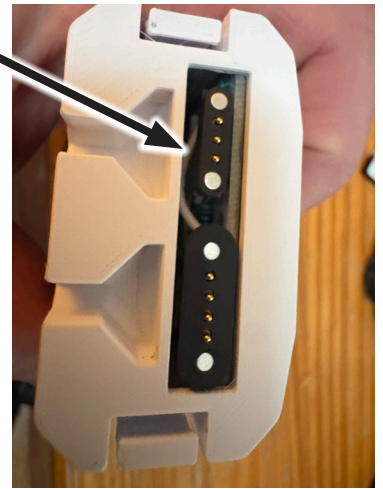


```
Time and Place  rtc!=gps
2005-04-12 00:27:56 UTC
Set clock on serial link
GPS fix: False  #Sats: None
latitude: None deg
longitude: None deg
altitude: None m
RETURN
```

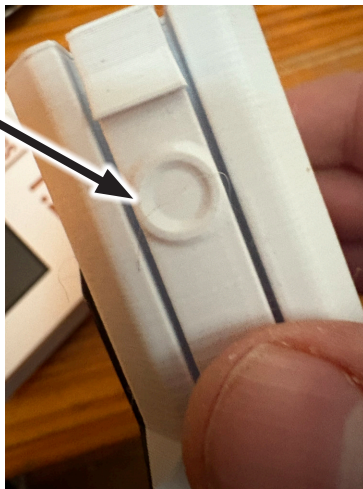

Remote Sensing (RS) Plugin



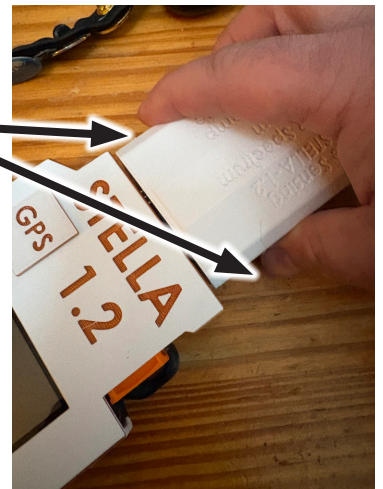
Magnetic Connectors



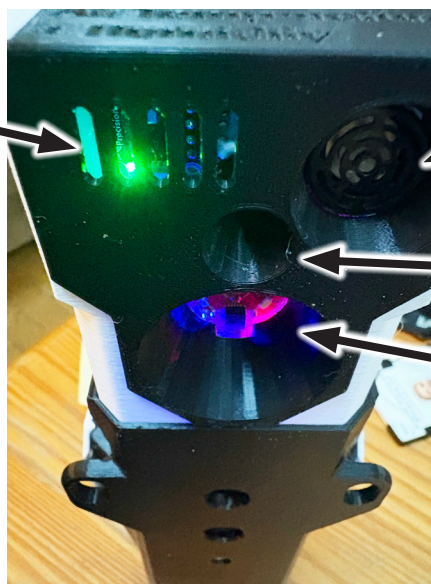
Plugin Clip



Press Both Clips Simultaneously On Each Side to Snap In/ Take Out Plugin



Temperature & Humidity



Ultrasonic Rangefinder

Surface Temperature

18-band Spectrometer

Tips: Do not point spectrometer towards the sun. When taking measurements stay out of the instrument and your own shadow. The spectrometer angular field of view is 40° so the height to swath diameter is approximately 1:1 - at 1 meter high the swath is about 1 meter in diameter.

Light

Remote Sensing Plugin Required

Y-Axis Options:

- Counts
- Normal Counts (normalized)
- Irradiance (W/cm²)

Scale:

- Linear or Log

Spectrum Selection

- Visible only
- NIR only
- VIS + NIR (default)
- UV options if sensor attached

Spectral Plot

- Bar colors match wavelength bands
- NIR bands shown in black
- Bar widths proportional to bandwidth
- Educational: shows measured regions

X-Axis:

- 400-1000nm range

X-Axis Unit Selection:

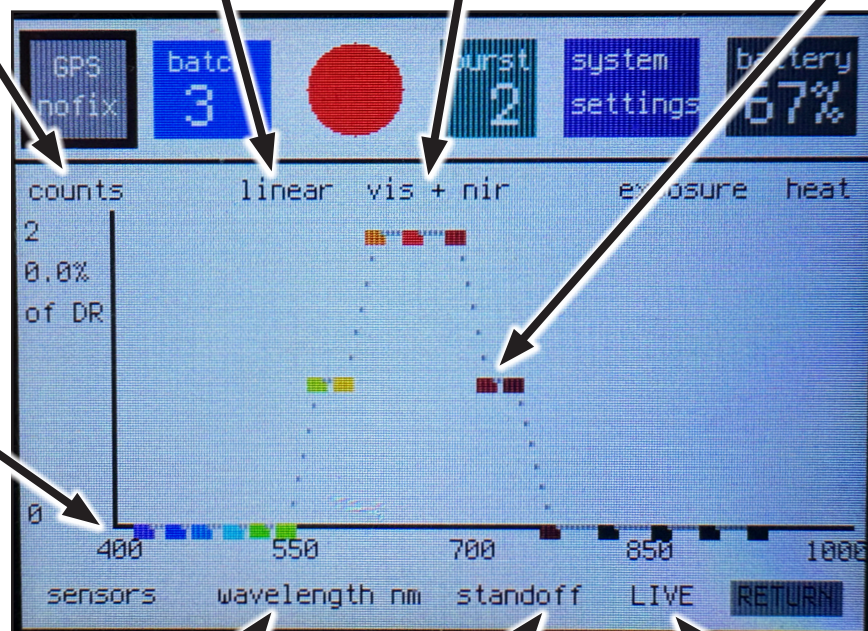
- Wavelength nm (default)
- Energy eV (electron Volts)
- Frequency THz (Terahertz)
- Wavenumber / cm

Standoff:

Shows distance 0.3-2.5m

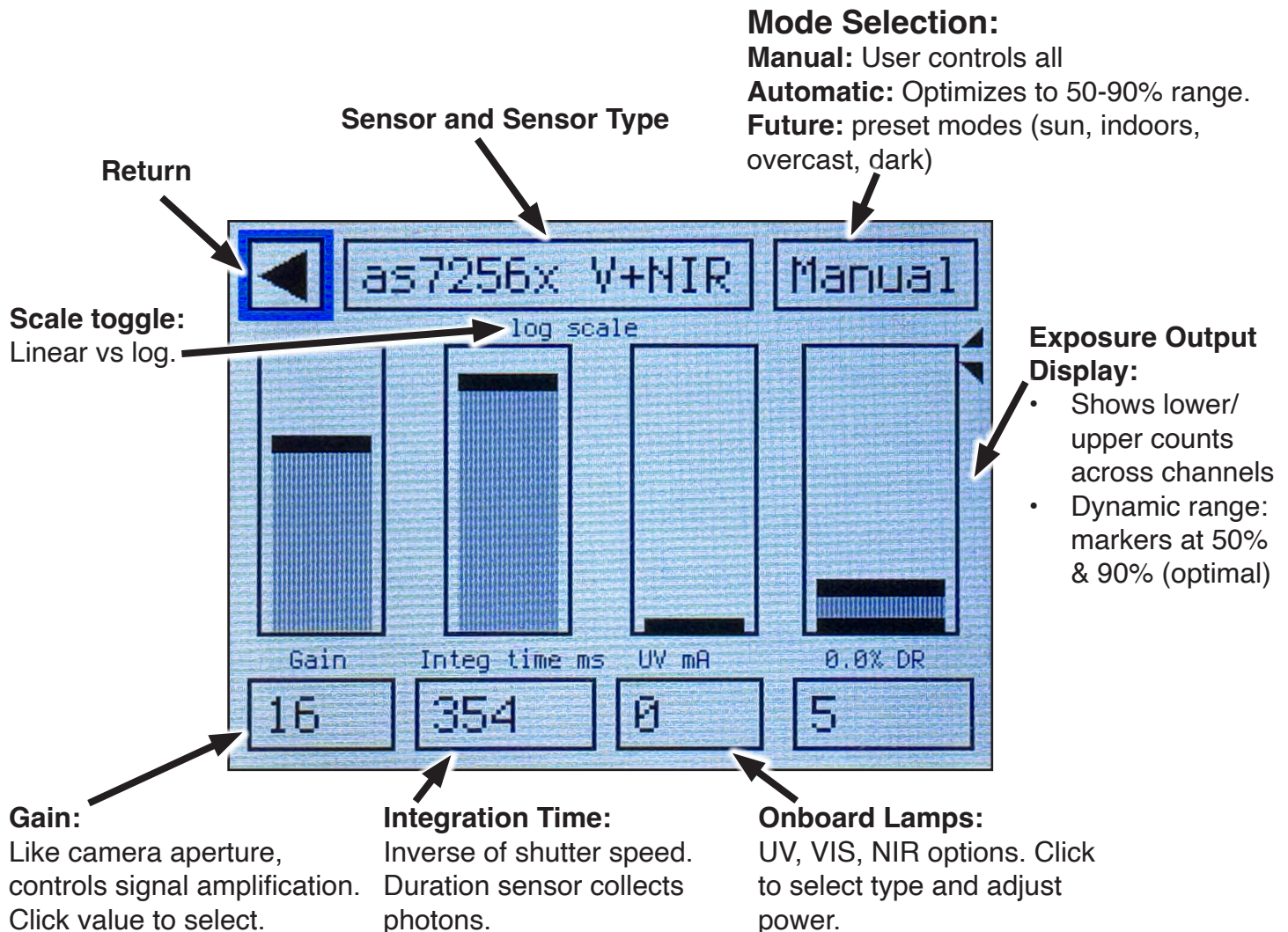
Live/Hold:

Update or freeze plot



Exposure

Remote Sensing Plugin Required

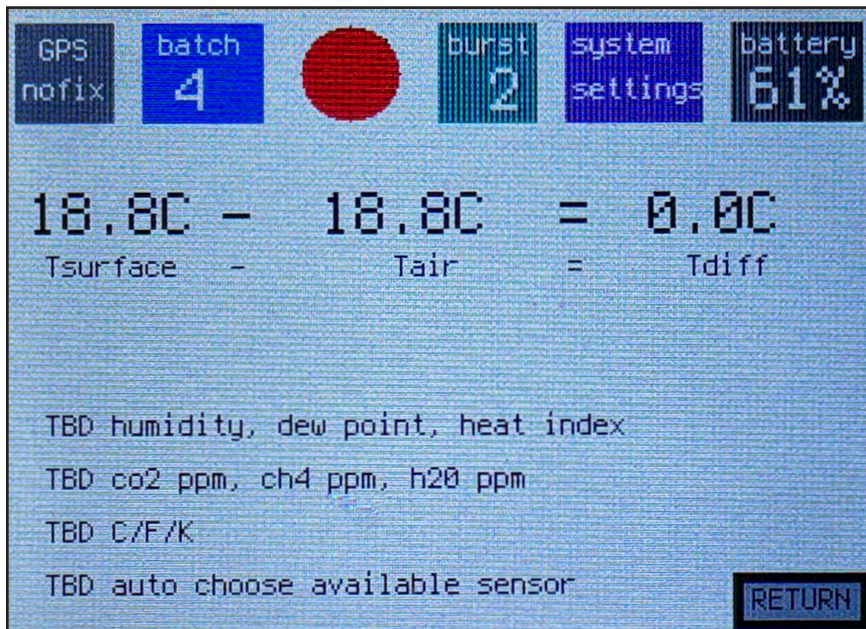


Exposure: (example)

When measuring a plant leaf indoors, you point STELLA at the target and notice the exposure bar shows only 5% of dynamic range—far too low for quality data. You click on the gain value, currently set to “16,” and select “64” from the dropdown to quadruple the signal amplification, then click the integration time and change it from “50ms” to “200ms” to allow more photon collection time. Watching the right side display, you see the bar graph rise as your adjustments take effect, and the highest count value climbs from barely visible to a healthy position between the 50% and 90% triangle markers—now you’re in the optimal range for accurate measurements with good signal-to-noise ratio and no saturation.

Heat

Remote Sensing Plugin Required



Surface Temperature Sensor:

- Measures thermal IR radiation
- Assumes emissivity = 1 (matte surfaces)

Temperature Difference:

- Shows Surface - Air
- Difference = 0: Same temp
- Positive: Surface warmer
- Negative: Surface Cooler

STELLA-1.2 RS Plugin Use Procedure

Remote Sensing Plugin Required

Calculate Required STELLA Height Above Target:

STELLA height is calculated to ensure only the target is being measured and there are no foreign objects recorded to influence radiance/reflectance.

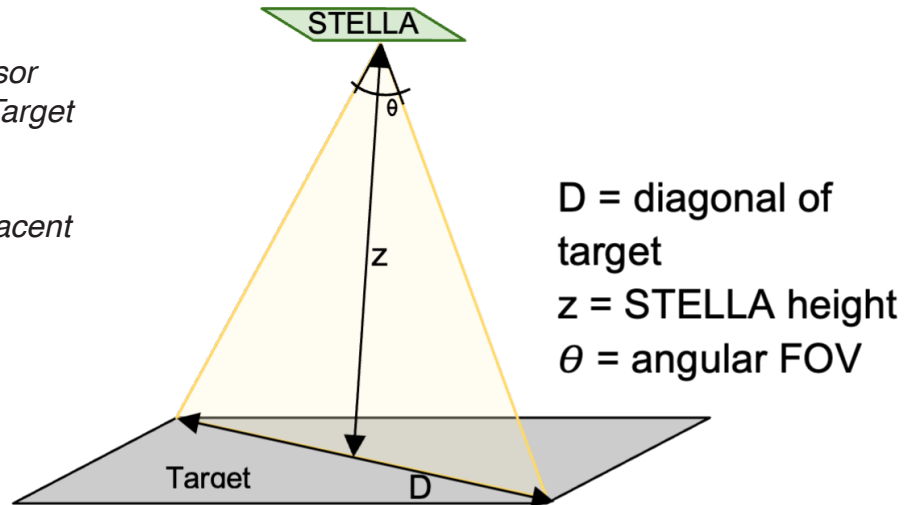
**STELLA-1.2 has an Ultrasonic Rangefinder to collect height data*

Needed for Calculation:

- a. Angular FOV of Sensor
- b. Diagonal Length of Target

Equation:

$$\tan(\theta) = \text{Opposite} / \text{Adjacent}$$



a. The STELLA spectrometer uses the AS7265x Multi Spectral Chipset Sensor. The specs can be found here, including the angular FOV of 41° (Appendix. 1).

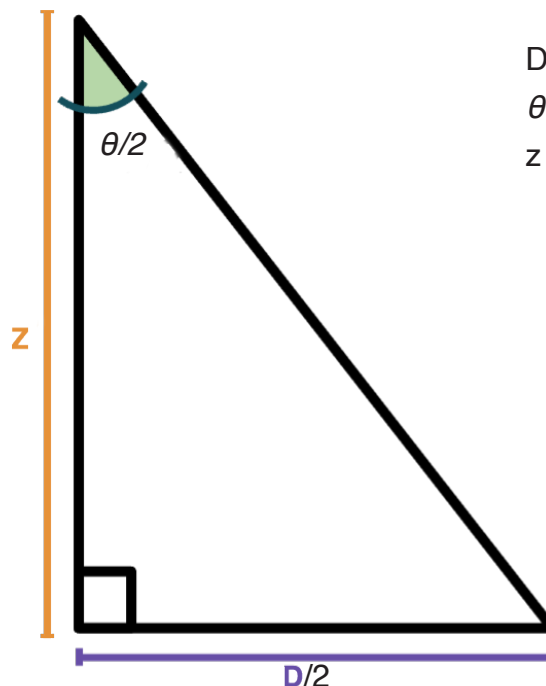
b. Target diagonal can either be measured or calculated using pythagorean.

Calculate the height using the equation below:

$$\tan(\theta) = \frac{\text{Opposite}}{\text{Adjacent}}$$

$$\tan(\theta / 2) = \frac{\frac{D}{2}}{Z}$$

$$Z = \frac{D}{2 \tan(\theta / 2)}$$



D = target diagonal
 θ = angular FOV
z = STELLA Height

Calculate STELLA FOV:

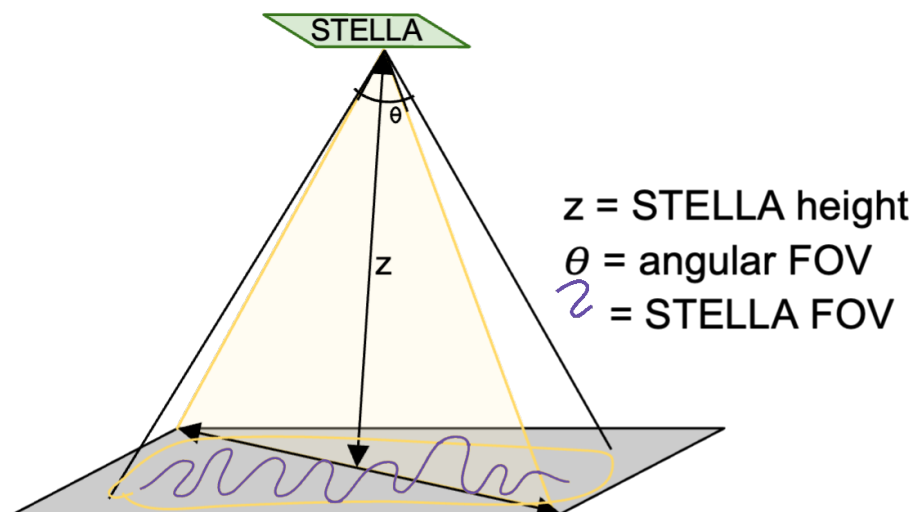
STELLA FOV in area will be used to convert measured irradiance into reflectance, but it can also be used as a check to ensure the height is correct.

Needed for Calculation:

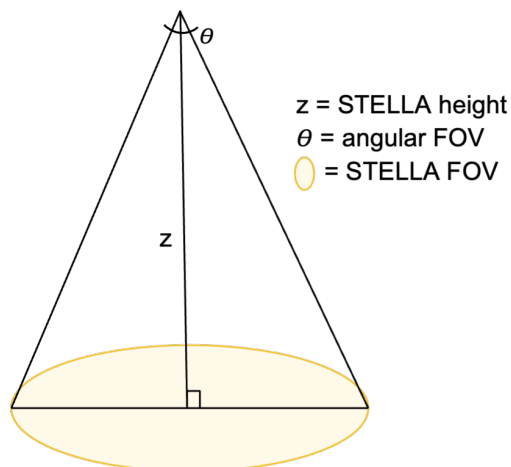
- STELLA Height
- Angular FOV of Sensor

Equation:

$$A = \frac{\pi}{4} \cdot D^2$$

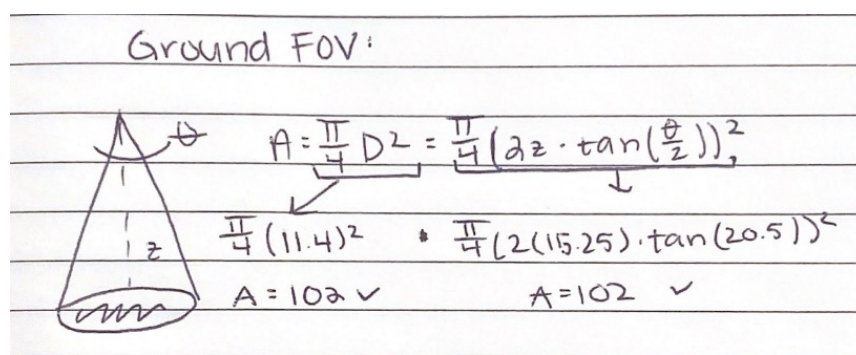


Using the angular FOV and height from the previous step, the STELLA FOV on the ground can be calculated using the equation below:



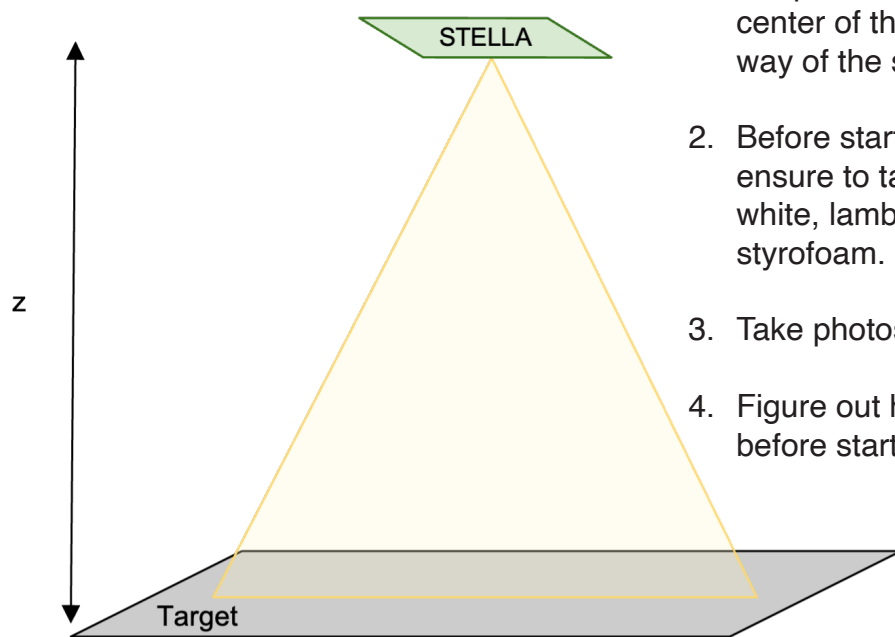
$$A = \frac{\pi}{4} \cdot D^2$$

$$A = \frac{\pi}{4} \cdot (2z \cdot (\frac{\pi}{4}))^2$$

Example:

Collect Measurements:

Following a consistent method/technique for collecting measurements using the STELLA can reduce error. It is also important to take notes during collection to document conditions to look back on as potential causes for outliers.



STELLA spectrometer collection methods

General Good Practice:

1. Keep the STELLA at a consistent height above the center of the target. Be sure that you are not in the way of the sensor!
2. Before starting and between every new target, ensure to take a few Reference scans of a pure white, lambertian surface such as clean, white styrofoam.
3. Take photos of each target and set up.
4. Figure out how many measurements are needed before starting-- lots of duplicates!

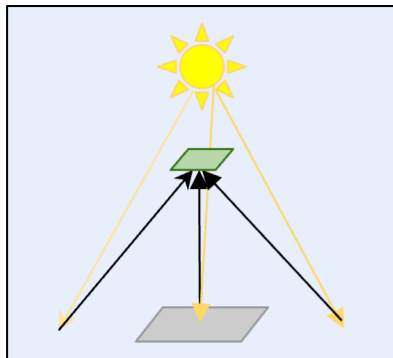
STELLA Collection Records Template:

| | |
|---|------------------------------------|
| Date: 7/25/24 | STELLA: <u>2544</u> tool ID |
| Notes: 17°C, slightly cloudy no rain, sunlight, wind speed: 6 mph NW, measurements taken facing N, 38cm above target conditions, i.e., weather, lighting, temperature, etc | |

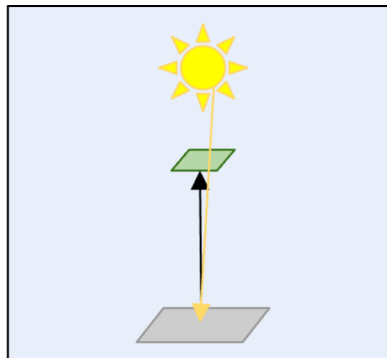
| Time: | Batch #: | Description: | Time: | Batch #: | Description: |
|---------|----------|----------------|--------------------|---------------|-----------------------------------|
| 10:42am | 31 | REF | 10:44am | 36 | Moss Green Gel |
| 10:42am | 32 | REF | 10:44am | 37 | Moss Green Gel mistake |
| 10:43am | 33 | Moss Green Gel | 10:45am | 38 | Moss Green Gel |
| 10:43am | 34 | Moss Green Gel | 10:46am | 39-43 | REF |
| 10:43am | 35 | Moss Green Gel | 10:48am | 44 | Primary Red Gel |

Convert from Irradiance to Reflectance :

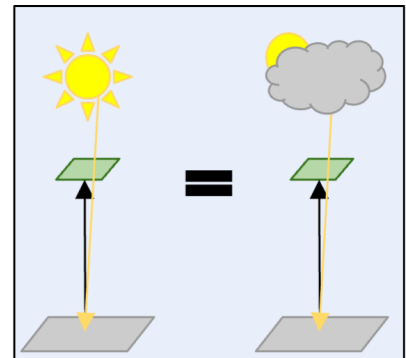
STELLA measurements are recorded in irradiance, but it needs to be converted to reflectance for standardization. Reflectance accounts for area of the target and variation from differences in illumination conditions.

**Irradiance**

Measure of light off a surface per unit area (W/m²)

**Radiance**

Measure of light coming off a surface from a solid viewing angle (W/m²/Sr)

**Reflectance**

Standardized measure of light coming off a surface, consistent across illumination conditions (unitless)

Needed for Conversion:

- Irradiance Measurement of Target*
- Irradiance Measurement of White Reference*
- Ground FOV Area*
- Distance from STELLA to Target*

Equations:

*Irradiance to Radiance: irradiance * (distance²/Area)*

Radiance to Reflectance: target radiance / white reference radiance

Example:

Target Irradiance = 1149 uW/cm²
 Reference Irradiance = 20358 uW/cm²
 FOV Area = 102in² = 259cm²
 Height (Distance) = 15.25in = 38cm

$$\begin{aligned} \text{Target Radiance} &= \text{Target Irradiance} * \frac{\text{Height}^2}{\text{Area}} \\ &= 1149 * 38^2 / 259 \\ &= 6406 \end{aligned}$$

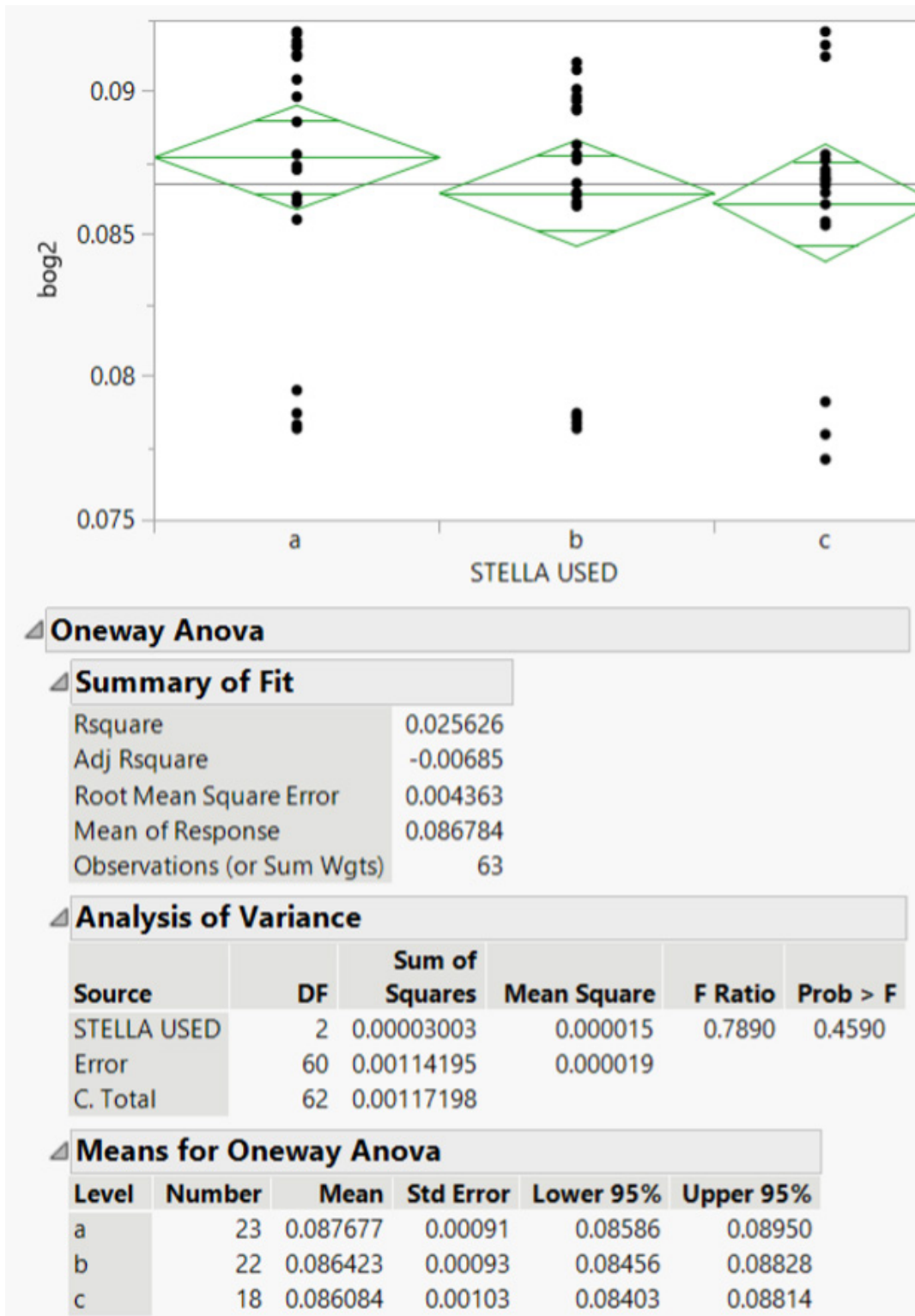
$$\begin{aligned} \text{Radiance} &= \frac{\text{Target Radiance}}{\text{Ref. Radiance}} \\ &= 6406 / 13501.7 \\ \text{Radiance} &= 0.056 \end{aligned}$$

$$\begin{aligned} \text{Reference Radiance} &= \text{Reference Irradiance} * \frac{\text{Height}^2}{\text{Area}} \\ &= 20358 * 38^2 / 259 \\ &= 13507.7 \end{aligned}$$

Data Processing and Statistics:

Running statistics on collected data will allow for determining the separation between measurements of the same target and see if there are any statistically significant differences between the samples.

ANOVA Example:



STELLA -1.2 Data Sample

**Measurement
Number:**

Instrument ID-Time
UTC-session-
session number

session number =
measurement for the
day

Instrument ID:
Unique ID
derived from
microprocessor

Timestamp:
Time in UTC
(Coordinated
Universal Time)

Decimal Hour: hours, minutes and seconds can be handled as a unified value

Batch Number: all measurements within continuous measurement collections and individual burst measurement

Burst Counter:
number of
measurements
taken per
recording

Integration time in milliseconds: the duration the detector collects photons to generate a single spectrum, acting similar to a camera's shutter speed

Irradiance Watts per Meter Squared: (spectral measurements)
quantifies the solar power incident on a surface per unit area

Sensor rows:
GPS Sensor:
(example)
divides the parameter
units and values into
understandable
measurements

| instrument_id | measurement_number | timestamp | decimal_hour | batch_number | burst_counter | sensor_name | part_number | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | value | parameter_units | 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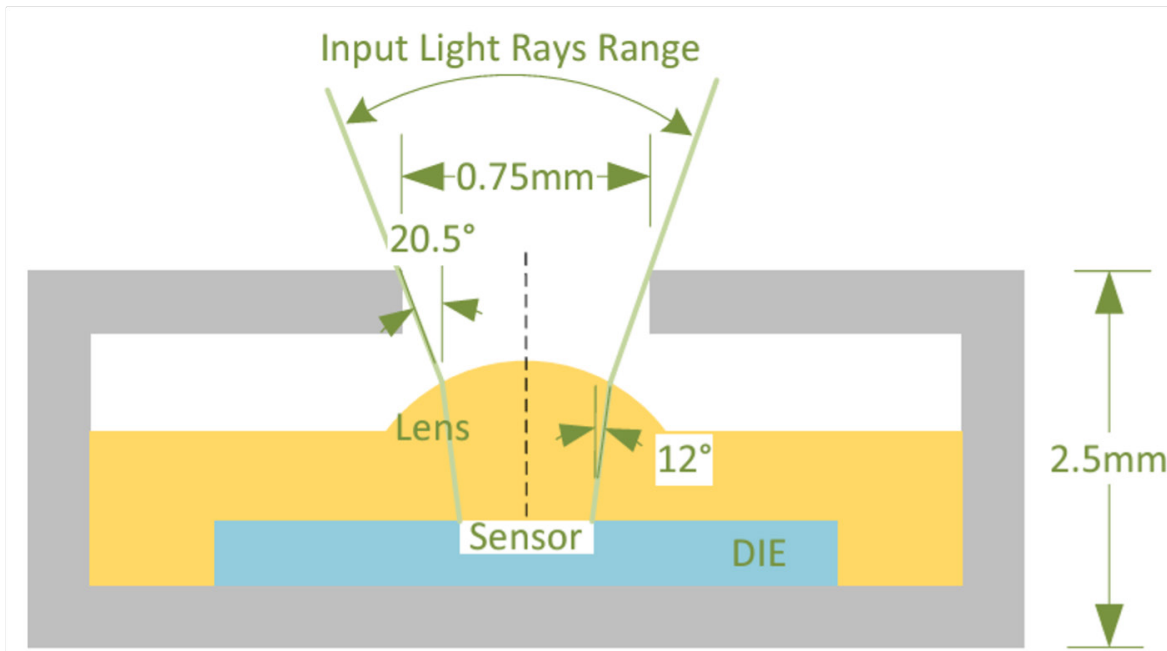
STELLA RS
18-bands:
this will repeat for
each measurement,
each are considered
a sensor

Wavelength in nanometers

Gain: the amplification factor applied to a detected signal (light intensity) by the detector's electronics, increasing its sensitivity to faint light

Raw Counts:
direct, unprocessed
signals (photons, ions, or
electrons) detected by an
instrument, representing
the raw data output before
normalization






Normal Count per Second:
the duration the detector collects photons to generate a single spectrum, acting similar to a camera's shutter speed



Appendix 1. AS7265x Multi Spectral Chipset Design Considerations, Optic Considerations

STELLA Instruments Across Spheres

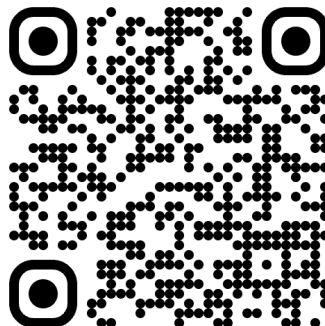
As a boundary-crossing complement to NASA's sphere-focused Earth science missions, STELLA multiplies their impact by creating the technical workforce that will design, build, calibrate, operate, and analyze data from the next generation of atmospheric, hydrospheric, biospheric, geospheric, and cryospheric observation satellites. Through participatory science, STELLA makes underlying mission principles accessible and achieves mission success through real-world, sphere-specific skill building.

| | | |
|---|---|---|
|  | GEOSPHERE Mission: GEMx, EMIT, NISAR, LAGEOS Application: Mineralogy Field Campaigns | STELLA Tech: STELLA-Q2 ✓ STELLA-1.2 RS + thermal ✓ STELLA-AQ ✓ Helio-STELLA ✓ Variables: Spectral curves for mineral ID, thermal properties, surface composition Cross-Sphere: STELLA-AQ ✓ for aerosol/dust correlation |
|  | HYDROSPHERE (Cross-sphere) Mission: Landsat, NISAR, SMAP, GPM, SWOT, ECOSTRESS, CYGNSS, Jason-3 Application: Soil Moisture & Agriculture | STELLA Tech: STELLA-SM ✗ STELLA-SM/R ✗ STELLA-DLS ✗ STELLA-1.2 RS ✓ STELLA-Q2 ✓ Variables: Soil moisture profiles, agricultural yield, vegetation stress, ET (beyond NDVI) Cross-Sphere: Integration with GPM precipitation validation |
|  | BIOSPHERE Mission: Landsat, PACE, MODIS, GEDI Application: Vegetation, Water, Surface Reflectance | STELLA Tech: STELLA-Q2 ✓ STELLA-1.2 RS + thermal ✓ STELLA Blue 1 ✗ STELLA Blue 2 ✗ STELLA DLS ✗ Variables: Leaving radiance, CDOM, water quality, phytoplankton/algae, HAB detection Cross-Sphere: Coastal interface connecting ocean-land |
|  | ATMOSPHERE Mission: TEMPO, Aura, Terra/Aqua, Suomi NPP, TROPICS Application: Air Quality Monitoring | STELLA Tech: STELLA-AQ ✓ Helio-STELLA ✓ STELLA Cloud Base ✗ STELLA-AM ✓ Variables: PM2.5/PM10, CO ₂ , aerosol optical depth, PBL dynamics Cross-Sphere: Mobile AERONET simulation (Helio + AQ) |
|  | CRYOSPHERE Mission: Landsat, ICESat-2, NISAR, SMAP, GRACE - FO, PREFIRE, ECOSTRESS, GPM Application: Snow/Ice Properties & Permafrost | STELLA Tech: STELLA-Q2 ✓ Helio-STELLA ✓ STELLA-SM/R ✗ + sensors 💡 Variables: Albedo, freeze-thaw, active layer thickness, permafrost carbon release Cross-Sphere: GPM precipitation phase (rain vs snow) |

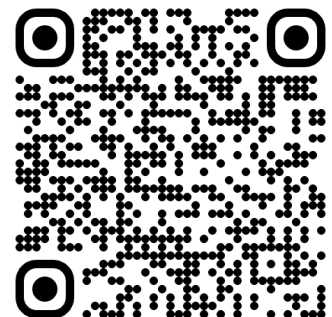
NASA Earth
Science Spheres



STELLA Website



STELLA Community Forum



Let's Build STELLA together!

Credit

Mike Taylor - Layout, Author
 Bianca Cilento - Layout, Author
 Paul Mirel - Author